

# CALIPSO and ATHENA-OAWL Performance Comparisons

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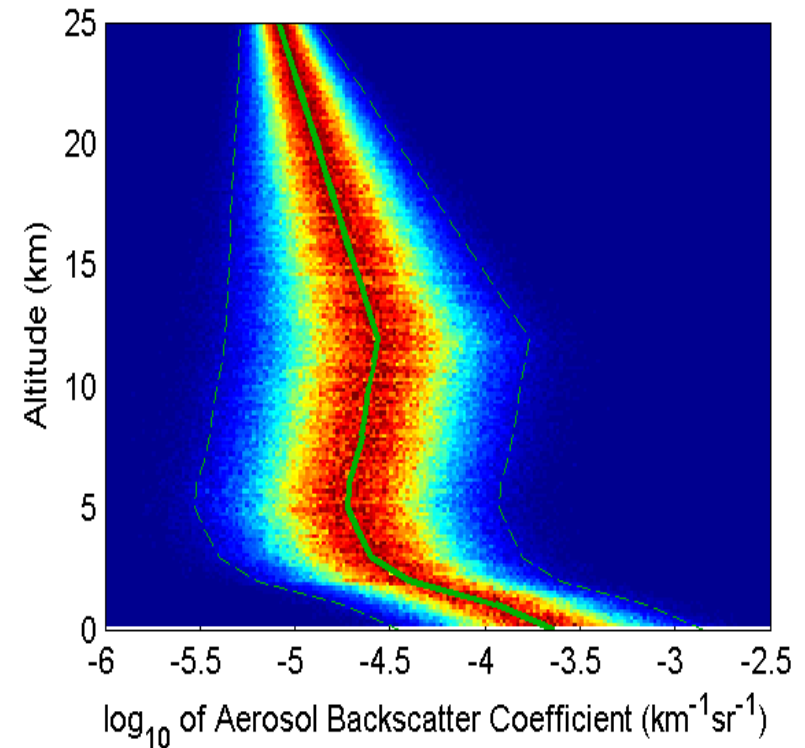
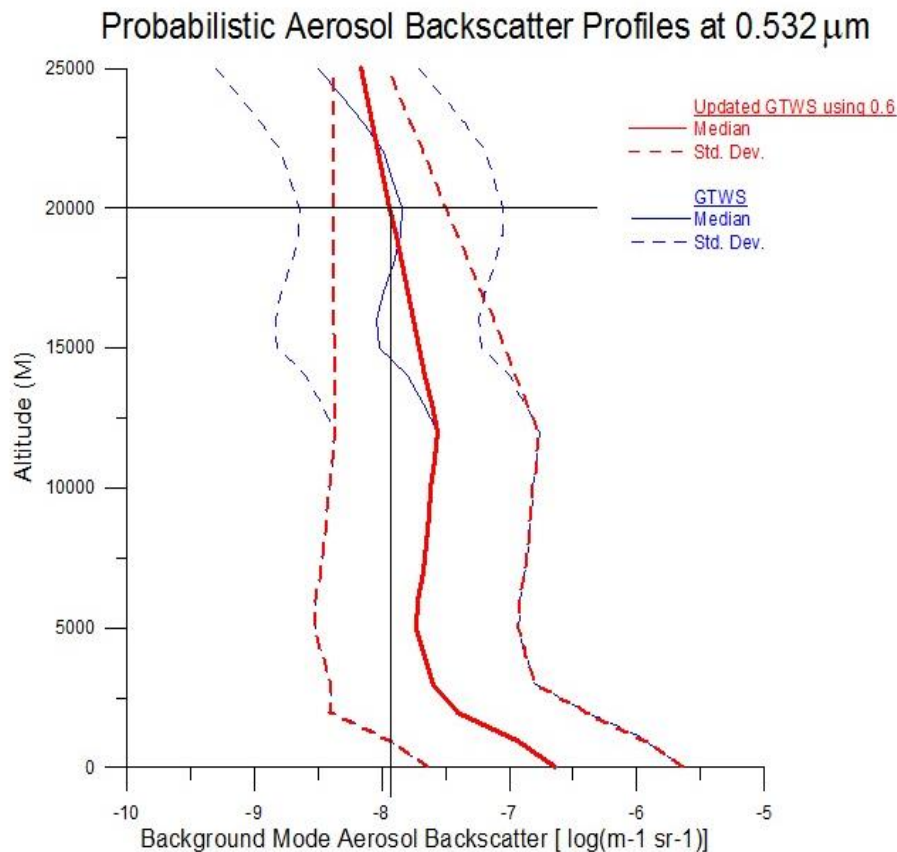
# Outline

- Predicted ATHENA-OAWL Performance
- SNR vs CNR definitions
- ATHENA-OAWL vs. CALIPSO
- CALIPSO sensitivity
- CALIPSO + SWA-GRABOP profiles
- Implications for ATHENA-OAWL



# GRABOP Distributions - SWA

Global Reference Atmospheric Backscatter Opportunity Profile  
(Dave Emmitt, SWA, Previous Talk)

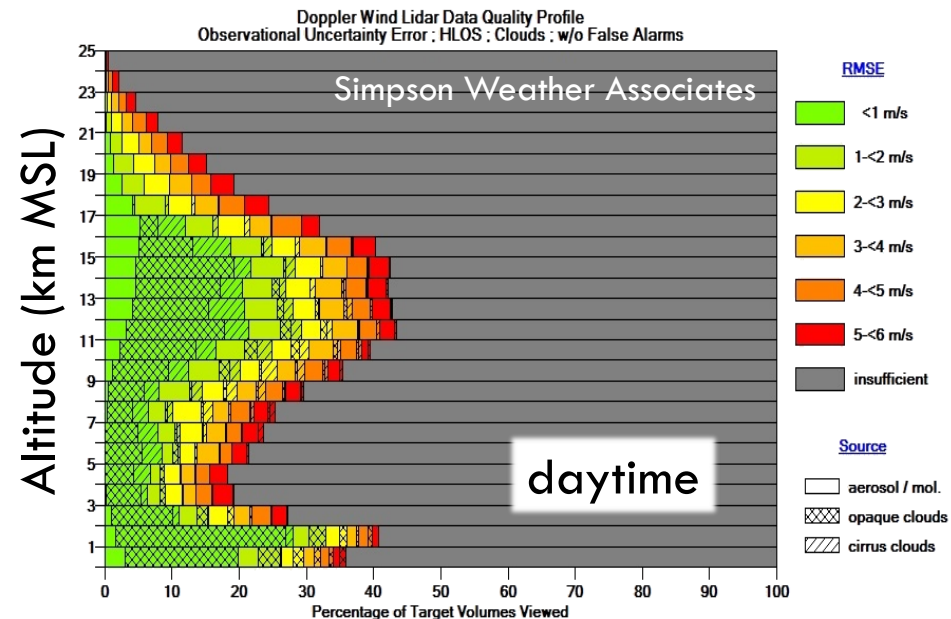
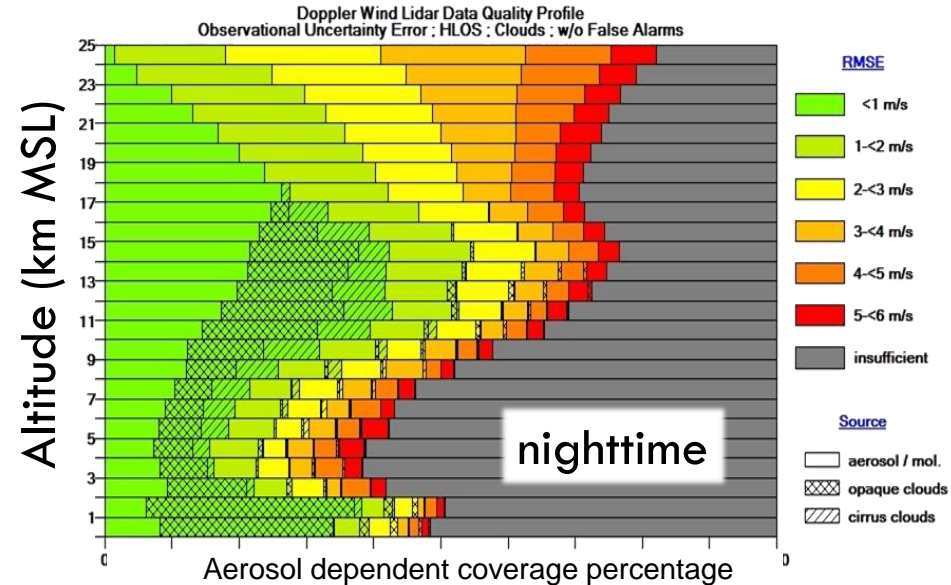


Altitude dependent distributions of aerosol backscatter coefficients based on the 532 nm GRABOP profiles.

# 532 nm ATHENA-OAWL

## Performance Prediction

- Performance Prediction Profiles from SWA
- Performance largely depends on aerosol loading
- 15.5 orbits → 8640 profile attempts/day (10s accum.)
- 2100-4000+ profiles (alt. dependent) with  $\leq 3$  m/s precision
- Compare to  $\sim 1200$  radiosonde profiles/day



# SNR vs CNR

$$P_r(R_z) = \frac{E_t \pi D^2}{R_z^2 \tau \sqrt{N}} T^2(R_z) \beta_{a+m}(R_z) \eta_t \eta_r$$

- CALIPSO “Signal-to-noise” (SNR): total return signal (from molecules and aerosols) vs. noise sources (signal shot, background, and dark)
  - ▣ CALIPSO users often incorporate “N” into SNR –
    - Not a per-shot SNR (ambiguous): “better SNR” achieved through more accumulation.
    - Signal *detectability* increases with  $N^{-1/2}$ .
    - But the terminology has become standard, so...
  - ▣ We clarify by using SNR(N) here.
- OAWL lidar signal will also use SNR(N)
- OAWL Wind “*Carrier signal to noise*” (CNR): related to fringe contrast, from the portion of the signal that forms a fringe in the interferometer
  - ▣ Wind speed precision improves with N (not CNR)



# ATHENA-OAWL vs. CALIPSO

Primary lidar equation parameters:

Instrument Parameter	CALIPSO	A-OAWL	SNR Ratios
Wavelength(s)	1064 nm and 532 nm	532 nm	-
Laser Energy @ 532 nm	110 mJ (per wavelength – 220 mJ total )	160 mJ	1.21 ( $P^{1/2}$ )
Laser PRF	20 Hz	150 Hz	2.74 ( $N^{1/2}$ )
Telescope Aperture Diameter	1 m	0.7 m	0.49 (Area)
Background Light Filtering	30 pm etalon filter and 1 nm interference filter	Same	-
Detector Type	PMT, QE 0.15	PMT, QE 0.18	1.1
Orbit Altitude	705 km	~400 km	3.11 ( $R^{-2}$ )
Pointing Angle	0.3° off nadir	40° off nadir	0.77

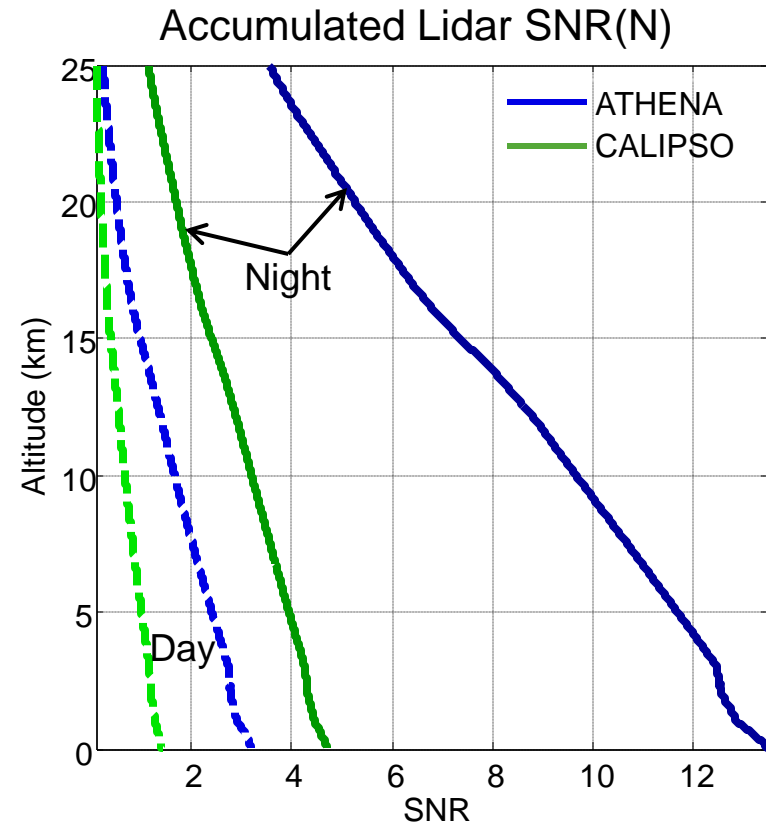
- For the same accumulation distances, range gates, altitudes, etc., ATHENA-OAWL will be  $>3x$  more sensitive than CALIPSO



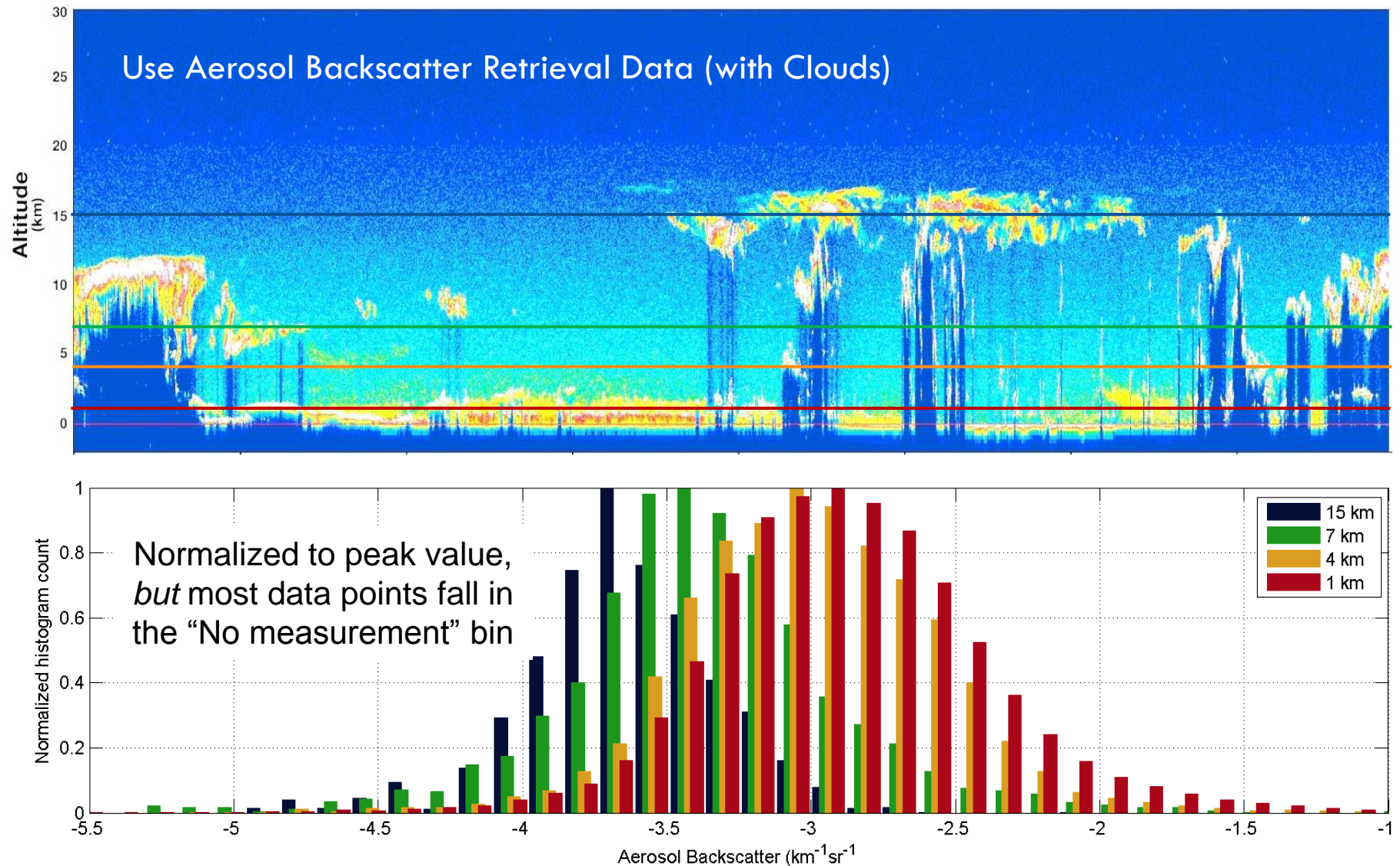


# 5km Lidar SNR(N) Comparison

- Radiometric model for OAWL based on that for CALIPSO
- ATHENA-OAWL & CALIPSO SNR
  - ▣ Day and Night performance
  - ▣ Same 5km ( $\sim 0.7$ s) accumulation
  - ▣ Same 30 m range gates
  - ▣ Uses the 532 nm GRABOP profile
- Different angles taken into account
  - ▣  $0.3^\circ$  for CALIPSO
  - ▣  $40^\circ$  for ATHENA
- 5km is for sensitivity comparison: OAWL will use 70-80km accumulation for winds
- 2X profiles from ATHENA (one per look angle)



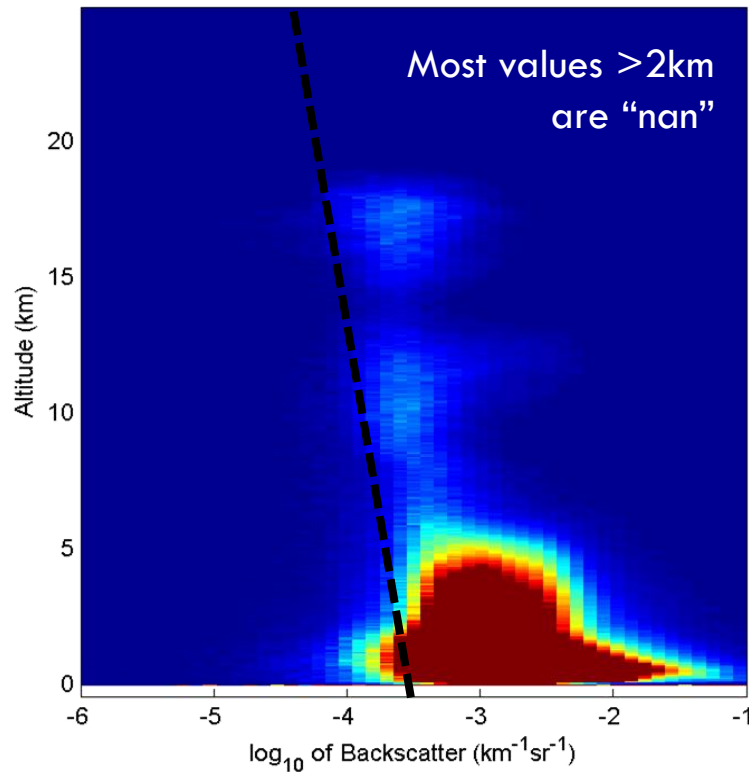
# CALIPSO backscatter: histograms from 5km profiles



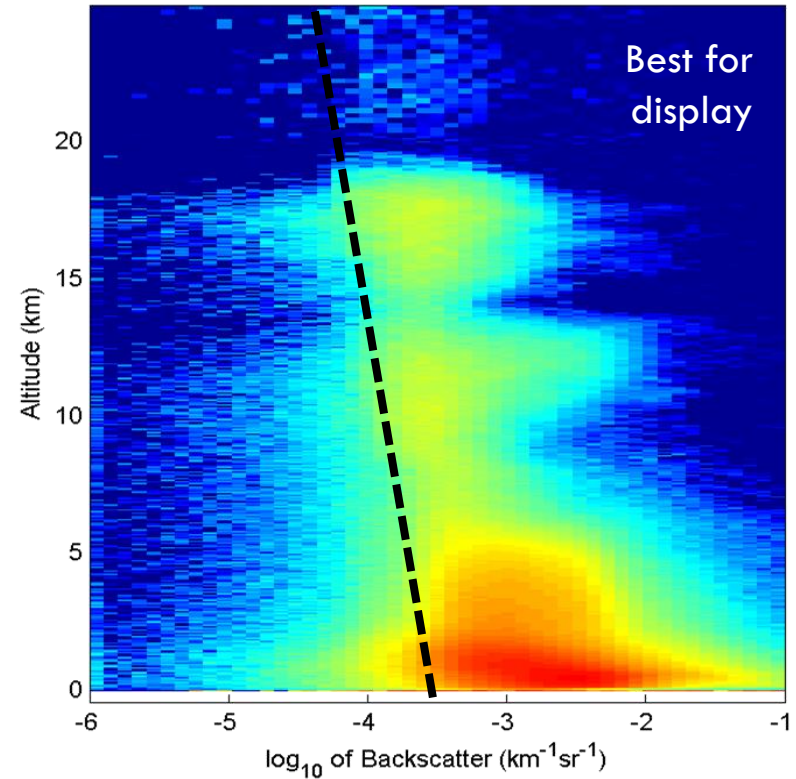


# Distribution for every altitude: in color

Altitude-relative distribution of  
all profiles (linear color)

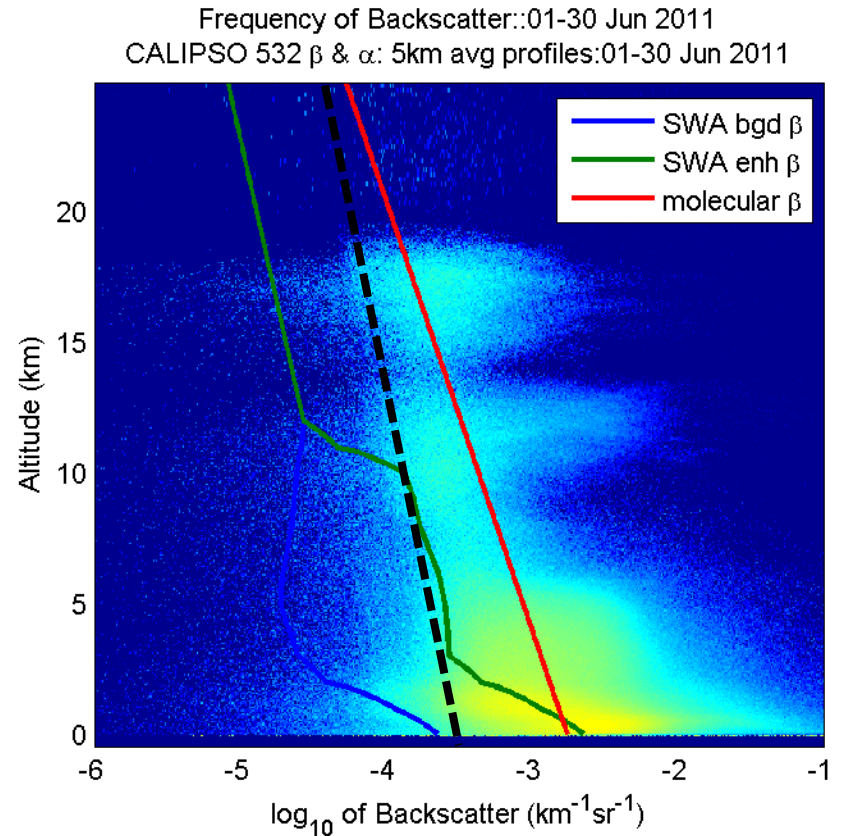


Altitude-relative distribution of  
all profiles (log-scale color)



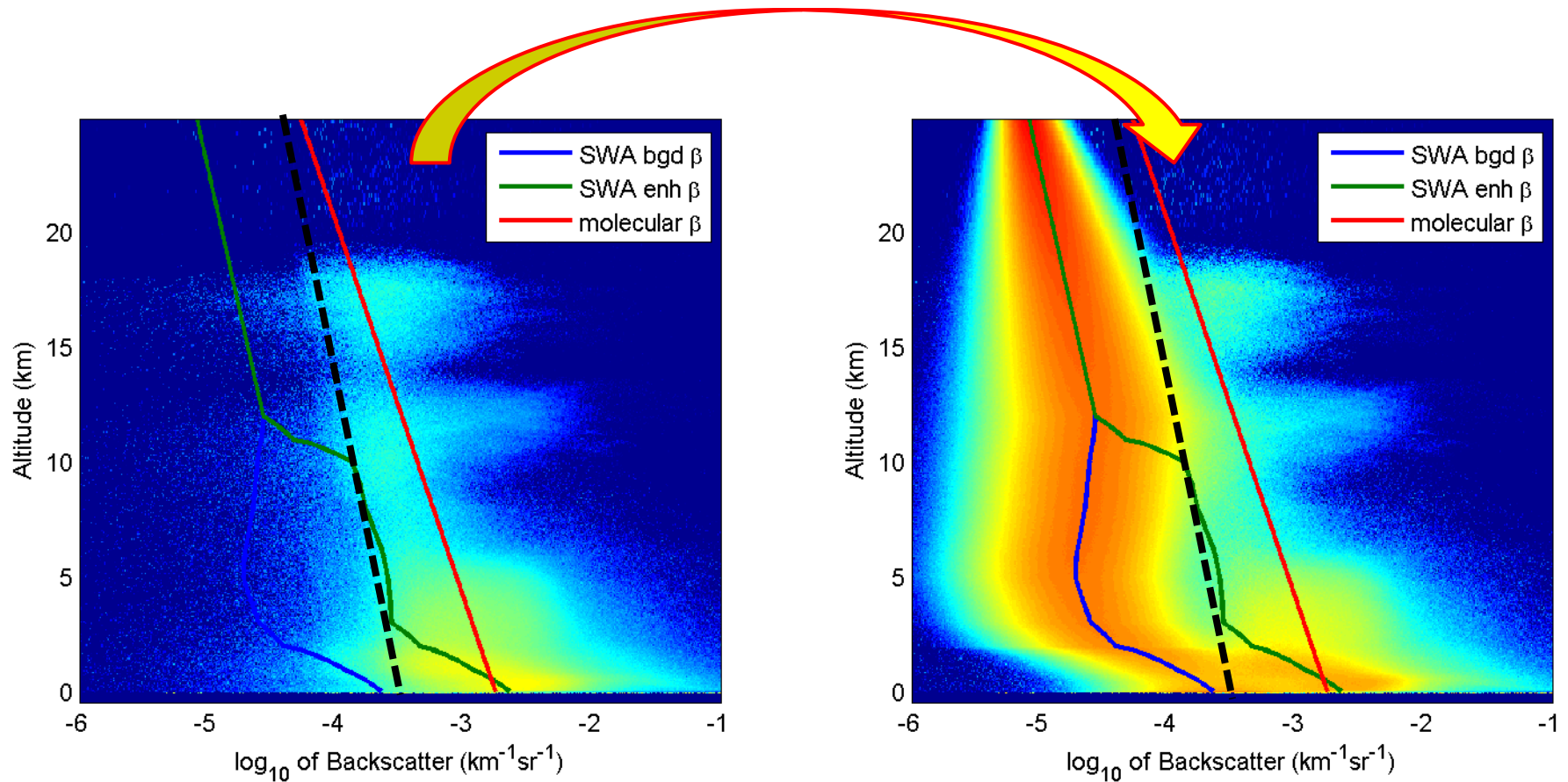
# CALIPSO Backscatter histograms vs. altitude

- 1 month (June 2011) of 5km CALIPSO profiles within  $\pm 56^\circ$  latitude
- Finer resolution (in  $\beta_a$  scale)
- Minimum CALIPSO detection vs. altitude
  - ▣ a sloped “edge”
  - ▣ edge smearing due to 5km averaging
- GRABOP profiles over-plotted for reference



# CALIPSO data + SWA Background Mode

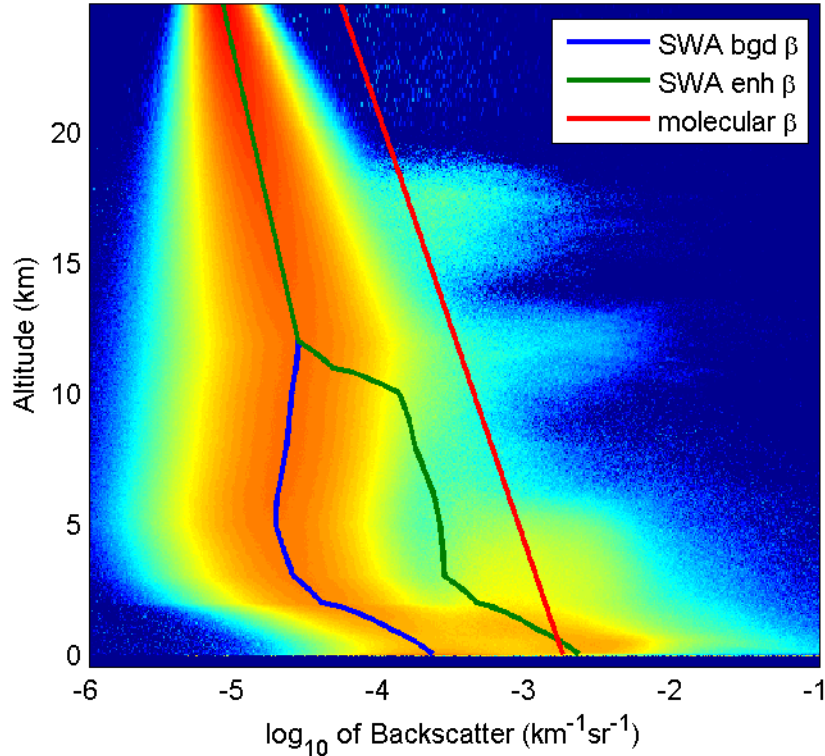
Fill in missing profile data with a random value from the GRABOP Distribution



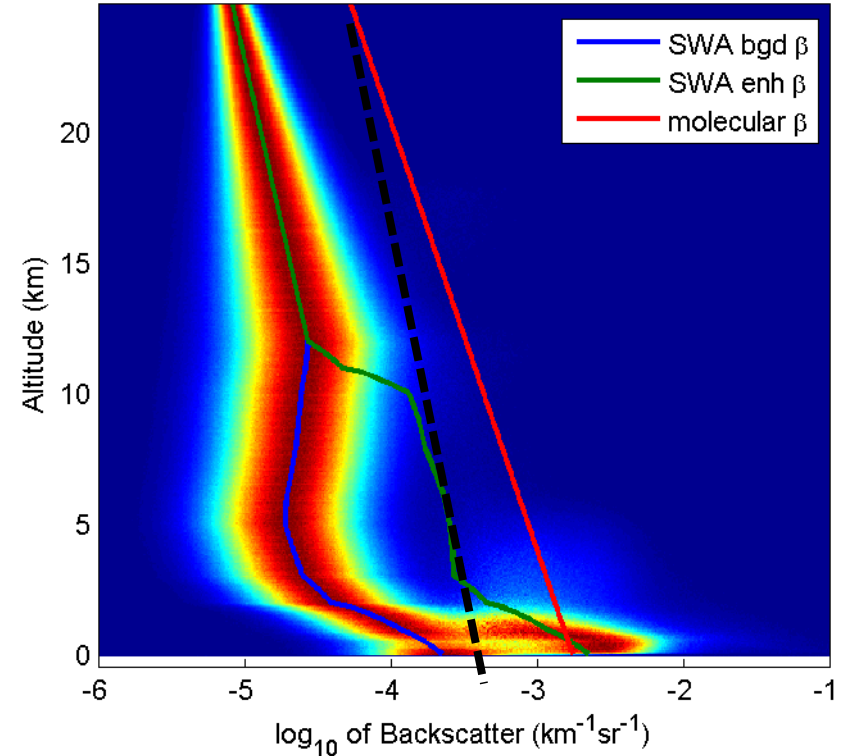
Log<sub>10</sub> histogram count color scale



# CALIPSO data + SWA Background Mode



Log<sub>10</sub> color of histogram count



Linear color of histogram count

# CALIPSO informs ATHENA-OAWL performance predictions

- CALIPSO doesn't see the background mode...
- ...but it tells us *where* the 532 nm background mode is conservative.
  - ▣ In the lower troposphere – 532 BGD is conservative
  - ▣ Higher altitudes – 532 BGD is likely a good basis for prediction
  - ▣ “True” distributions lie somewhere in between (altitude dependent)
- Performance prediction models also take into account aerosol and cloud **extinction** (Provided by CALIPSO, filled in with the background mode model values)

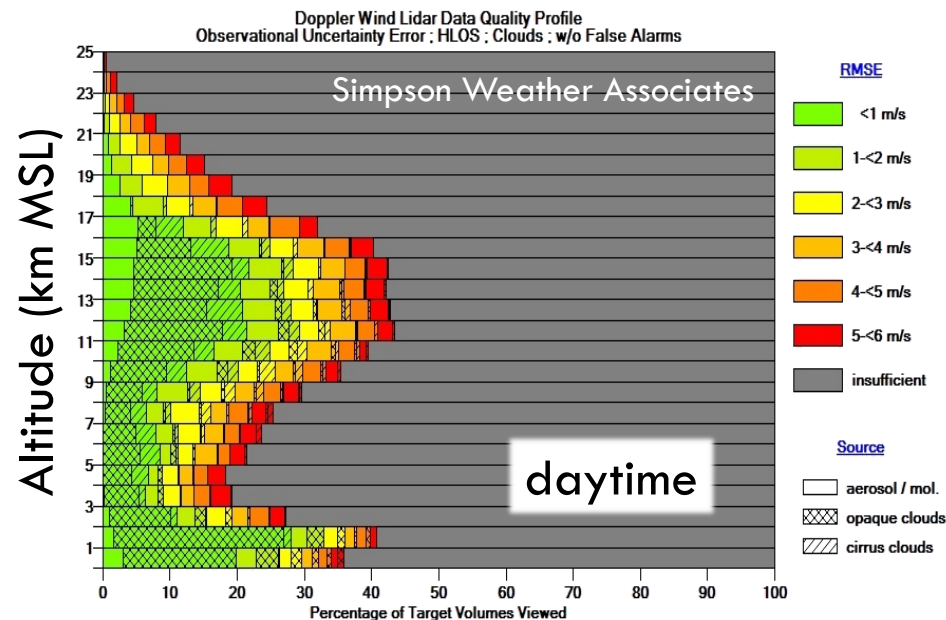
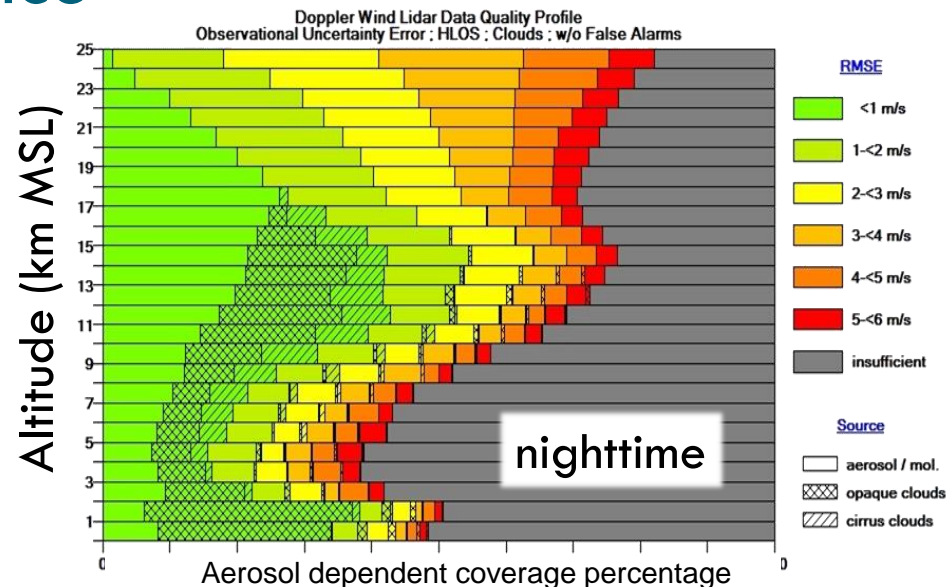




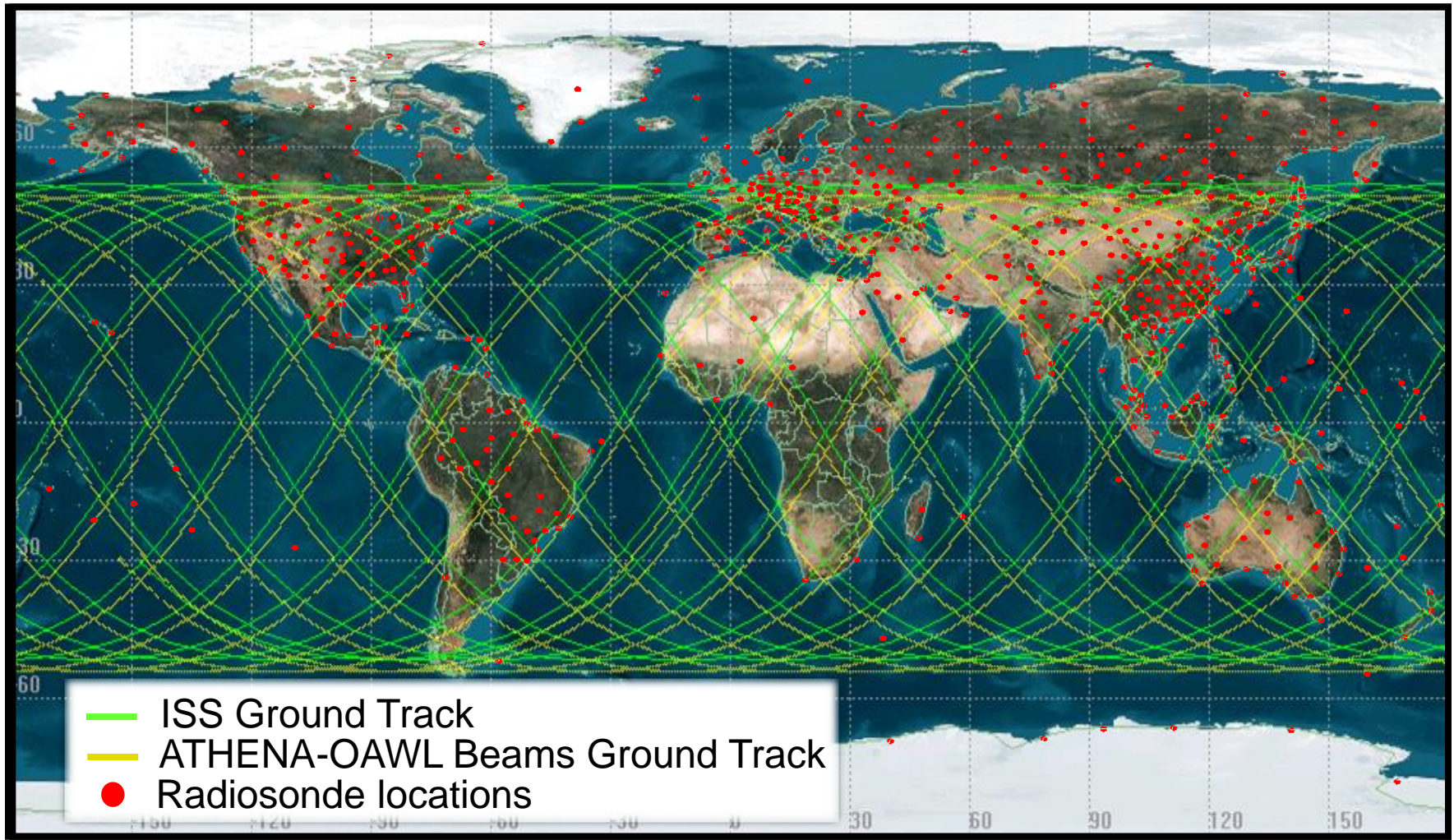
# 532 nm OAWL Performance

## Nighttime vs. Daytime

- Still based on GRABOP background mode only
- 8640 profile attempts/day (10s accum.)
- 2100-4000+ profiles per 24 hours (alt. dependent) with  $\leq 3$  m/s precision

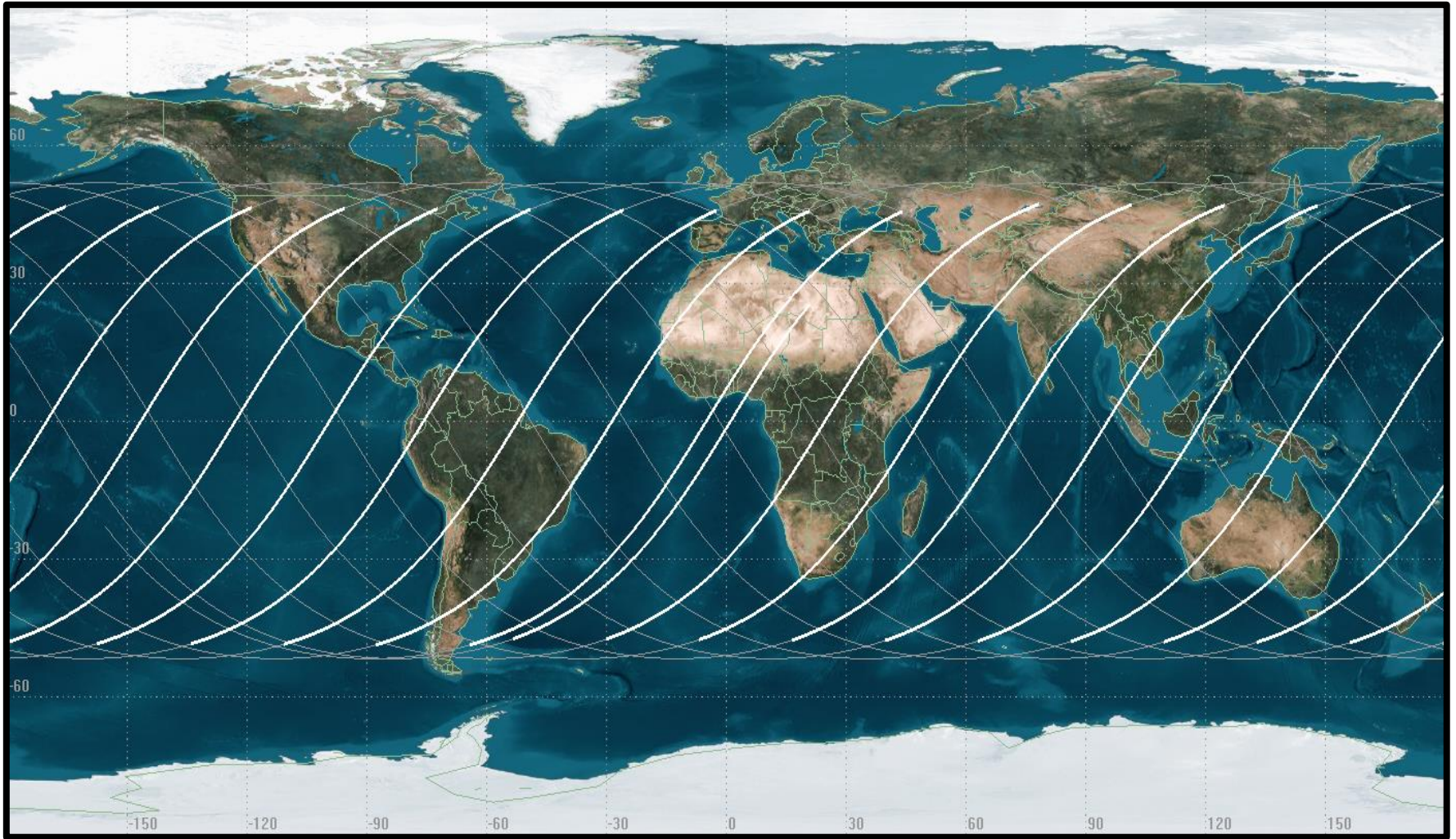


# OAWL ISS Coverage (Example 24 hours)

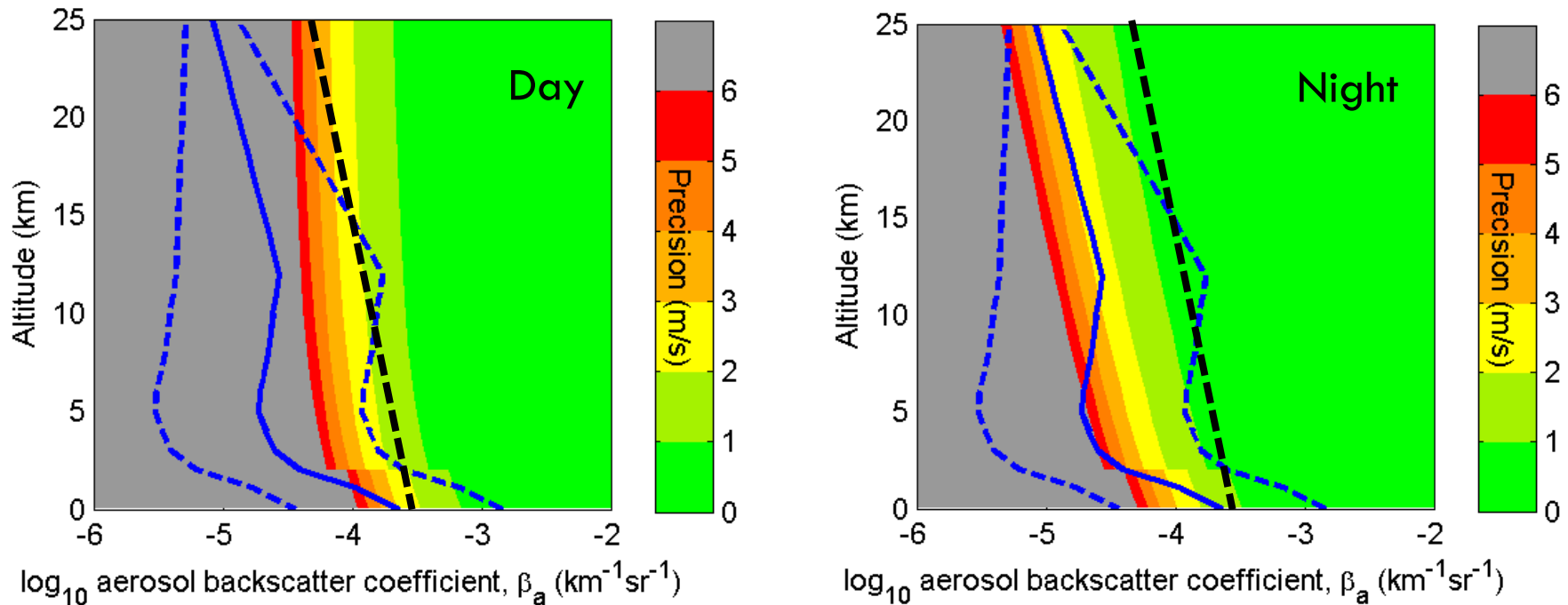




# ISS Tracks – Night vs. Day



# ATHENA-OAWL performance maps of precision vs. altitude/backscatter



Almost everywhere CALIPSO makes an aerosol backscatter measurement  
ATHENA-OAWL would have better than 3 m/s precision (day)  
and better than 2 m/s precision (night)

# Future efforts in this area...

- Generation of **shot coverage files** (timing, pointing, ISS motion, → maps of atmospheric coverage)
  - ▣ Strong STK (software good for orbit geometry) expertise at Ball
- Integration of **CALIPSO extinction** (including cloud extinction), filled in with GRABOP extinction profiles
- Extension from 1-month global distribution profiles to distribution profiles for **annual/spatial variations**
- Work with SWA to combine CALIPSO results, and CALIPSO/OAWL instrument models with **nature run** (clouds, etc.) to improve prediction estimates.
- Incorporation of **jitter/vibration/pointing** effects on performance (accuracy and precision)
- **Wind processing testbed:** Combine radiometric model with OAWL signal models and processing algorithms and use them to test/further develop:
  - ▣ pointing angle correction algorithms
  - ▣ cloud processing/weighting algorithms
  - ▣ Forward/aft overlap opportunities
  - ▣ LOS winds processed into horizontal wind speed & direction
  - ▣ Aerosol studies (two looks)





# Summary

- ❑ ATHENA-OAWL designs build on those from CALIPSO/CALIOP.
- ❑ CALIPSO performance helps to validate expectations for ATHENA.
- ❑ GRABOP background mode profiles indicate sufficient performance to meet ATHENA mission goals.
- ❑ 2x+ more profiles from ATHENA-OAWL than provided by the radiosonde network
  - ▣ With better global distribution
- ❑ CALIPSO data suggest there will often be more aerosols than predicted by background mode.
  - ▣ Good news for ATHENA-OAWL performance
- ❑ ATHENA-OAWL is expected to be at least 3x more sensitive than CALIPSO for the same accumulation time/range.
- ❑ Measurements in the forward-looking ATHENA-OAWL profiles will have overlap with the aft-looking profiles → 2x measurements for aerosol studies.

